

General Relativity

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Exercise Sheet 6

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Exercise 1: Motion in spherically symmetric gravitational field

We want to analyze the motion of freely falling objects in a spherically symmetric gravitational field $h_{\mu\nu} \equiv g_{\mu\nu} - \eta_{\mu\nu}$ of the form

$$h_{\mu\nu} = \begin{pmatrix} -2\phi(r) & 0 & 0 & 0 \\ 0 & -2\psi(r) & 0 & 0 \\ 0 & 0 & -2\psi(r) & 0 \\ 0 & 0 & 0 & -2\psi(r) \end{pmatrix}, \quad (1)$$

where ϕ and ψ are two a priori independent functions of the radial coordinate r .

1.1) Show that the geodesic equation is generally equivalent to

$$\frac{d}{d\tau}(\eta_{\mu\nu}u^\nu + h_{\mu\nu}u^\nu) = \frac{1}{2}\partial_\mu h_{\nu\rho}u^\nu u^\rho, \quad (2)$$

where $u^\mu = \frac{dx^\mu}{d\tau}$ is the 4-velocity with proper time τ .

Remark: In this equation $h_{\mu\nu}$ need not be small.

1.2) Let the curve be parameterized by $x^\mu = (t, x, y, z)$, so that $u^\mu = (\dot{t}, \dot{x}, \dot{y}, \dot{z})$, $\dot{} \equiv \frac{\partial}{\partial\tau}$. Evaluate the four equations (2) for (1) and prove the conservation law

$$\frac{d}{d\tau}((1 - 2\psi)(\dot{x}y - \dot{y}x)) = 0.$$

Assume next that the motion takes place in the $z = 0$ plane and replace (x, y) by polar coordinates (r, θ) . Show conservation of angular momentum:

$$L \equiv (1 - 2\psi)r^2\dot{\theta} = \text{const.}$$

1.3) Upon introducing $\rho \equiv \frac{1}{r}$ prove that the motion is governed by the equation

$$\left(\frac{K^2}{1 + 2\phi} - 1\right)\frac{1 - 2\psi}{L^2} = \left(\frac{d\rho}{d\theta}\right)^2 + \rho^2, \quad (3)$$

where K is an integration constant.

Hint: Recall the normalization $g_{\mu\nu}u^\mu u^\nu = -1$.

- 1.4) Perform an expansion of (3) in powers of ϕ . Show that the Newtonian limit is obtained to first order in ϕ for $\phi = \psi = -\frac{GM}{r}$, i.e., that this reproduces Kepler's elliptical orbits. Show that to next order, and taking the corrections implied by the Schwarzschild solution into account, the equations predict a precession of perihelia.